SCIMMA

Scalable Cyberinfrastructure for Multimessenger Astrophysics

0

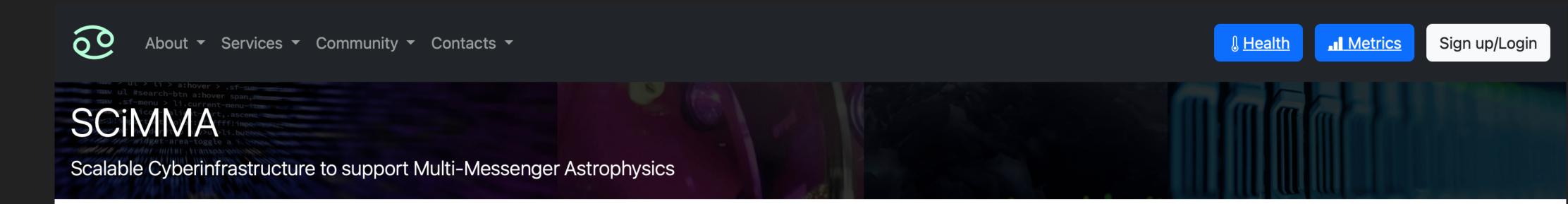
Andy Howell



University of California, Santa Barbara

+ SCIMMA, especially Jon Nation (LCO) & Curtis McCully (LCO)





NSF-sponsored group to study / build cyberinfrastructure for MMA, including stakeholders from LIGO, IceCube, computing centers, astronomers.

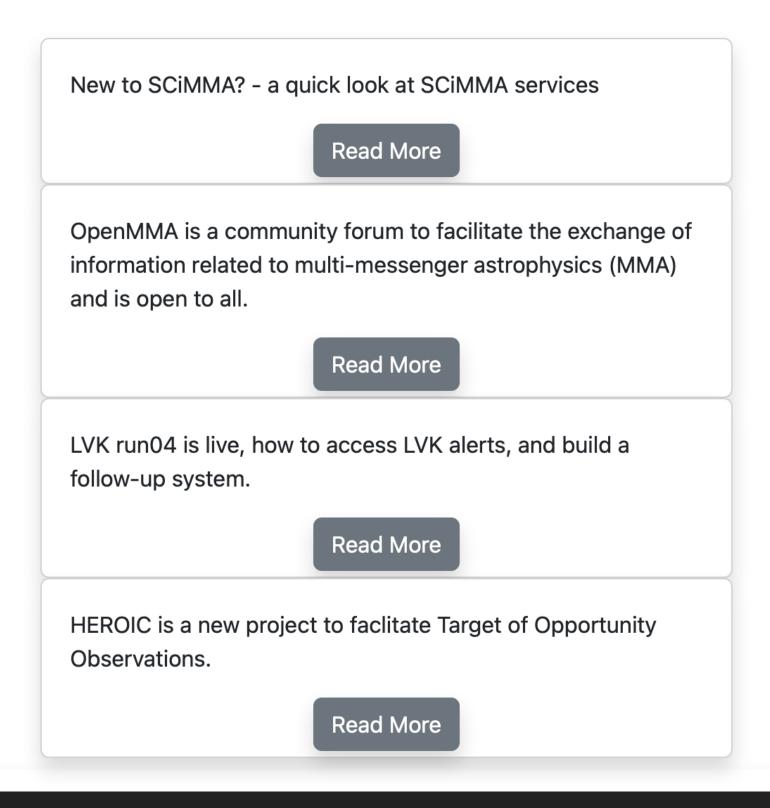
About SCIMMA

The goal of the Scalable Cyberinfrastructure to support Multi-Messenger Astrophysics (SCiMMA) project is to identify needed technical infrastructure through community engagement efforts, and then support these requirements by developing the required cyberinfrastructure and software systems. SCiMMA development projects draw on collaborations among astronomers, computer scientists, and data scientists.

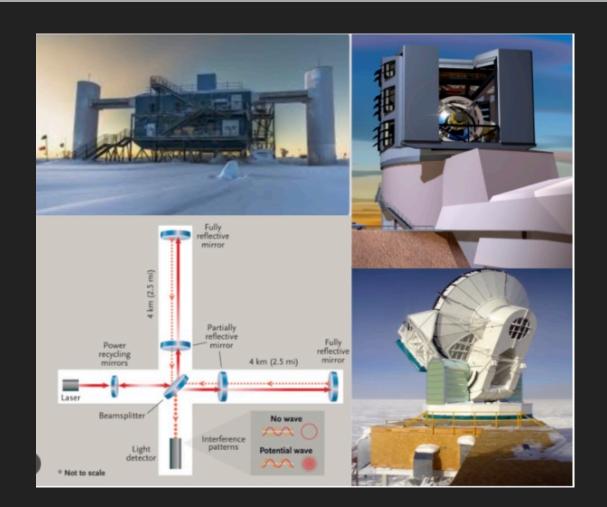
Multi-Messenger Astrophysics (MMA) signals are often transient, originating from distinct celestial events, making a rapid response to an MMA event a key capability, The response may include a fast analysis by a distributed science group; assembly of historical observations; determining the instruments available for new, prompt follow-up observations; instantating the observations; and resolving conflicting requirements for similar observations from multiple science groups.

SCiMMA is working to provide interoperating cyberinfrastructure to rapidly coordinate, combine, and analyze the large-scale distributed data from all these sources. The infrastructure must accommodate public communications (of interest to the community) and protected internal communications (to support work internal to a collaboration).

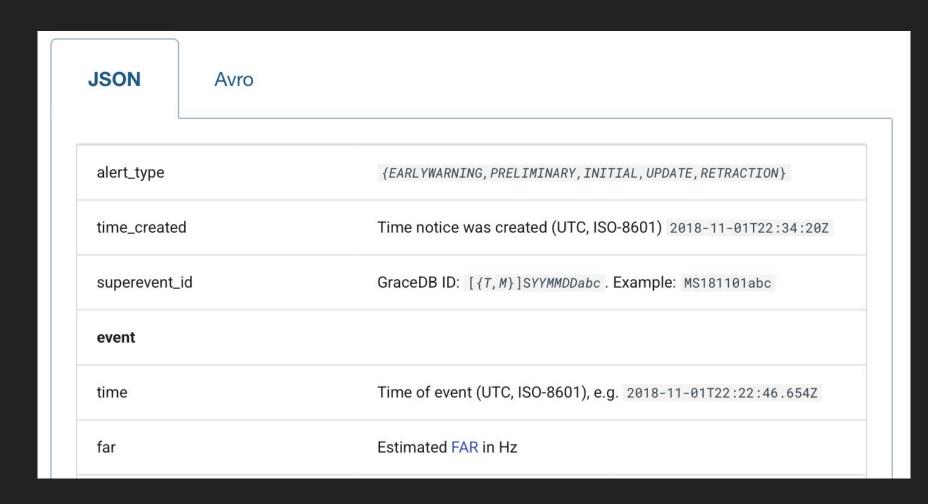
The collection of SCiMMA services increasingly statisfy the community's needs for Cyberinfrastructure. The OpenMMA forum provides an addition venue to assess progress and needs.



SCIMMA PRODUCTS



Open MMA
Community information



IGWN alerts
Gravitational wave alerts



Blast
Host galaxy information



Hopskotch/HERMES
Messaging



HEROIC
Telescope & instrument availability

BLAST

Blast - Q View Transients

Profile



A web service for the automatic, real-time characterization of transient host galaxies

Search the Blast database Q

Read the docs

Transient Processing Statistics

Blast: a Web Application for Characterizing the Host Galaxies of Astrophysical Transients

D. O. JONES ^{(1),*} P. McGill ^{(1),2,3,*} T. A. Manning ^{(1),4} A. Gagliano ^{(1),5,6,7} B. Wang (王冰洁) ^{(2),8,9,10} D. A. COULTER , 12 R. J. FOLEY , G. NARAYAN , 12, 13 V. A. VILLAR , 14 RAFF, A. W. ENGEL , 14 D. Farias , ¹⁵ Z. Lai , ¹⁶ K. Loertscher, ² J. Kutcka , ² S. Thorp , ¹⁷ and J. Vazquez ¹²

Blast • Q View Transients

△ Login

2025absa

Download data 🖖

Information

Transient

Name: AT2025absa

Right ascension: 9h29m15.21s Declination: -20d22m34.73s

Discovery Date: Oct. 26, 2025, 5:42 p.m.

Classification: None Redshift: None MW E(B-V): 0.046

Host

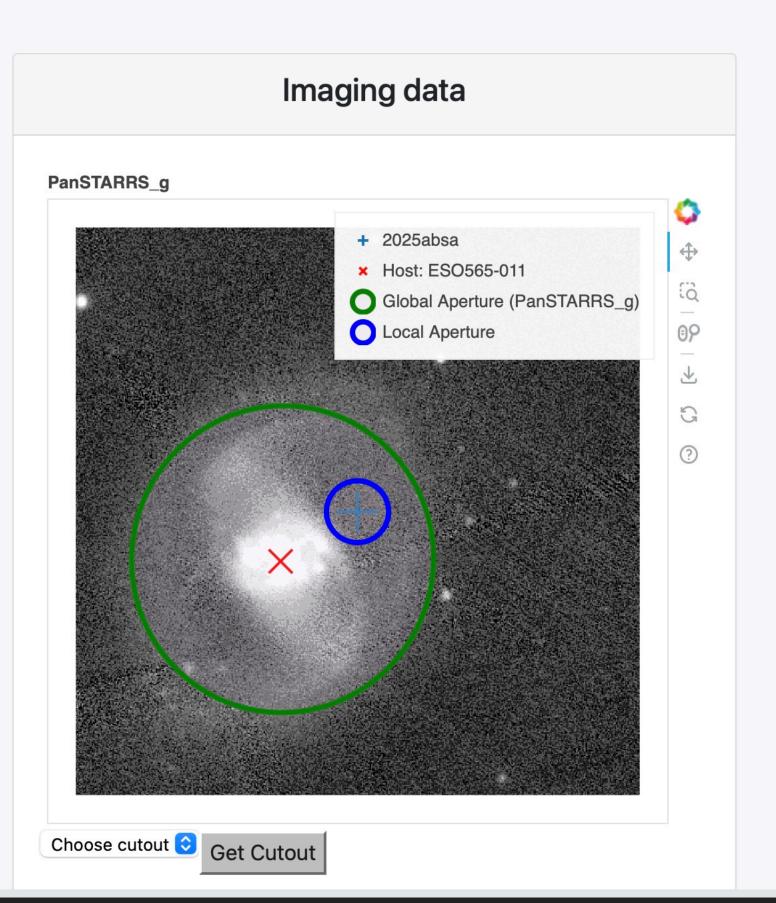
Name: ESO565-011

Right ascension: 9h29m16.42s Declination: -20d22m45.45s

Redshift: 0.014484 Photometric Redshift:

Processing Status

Cutout download



Aperture Photometry

Global

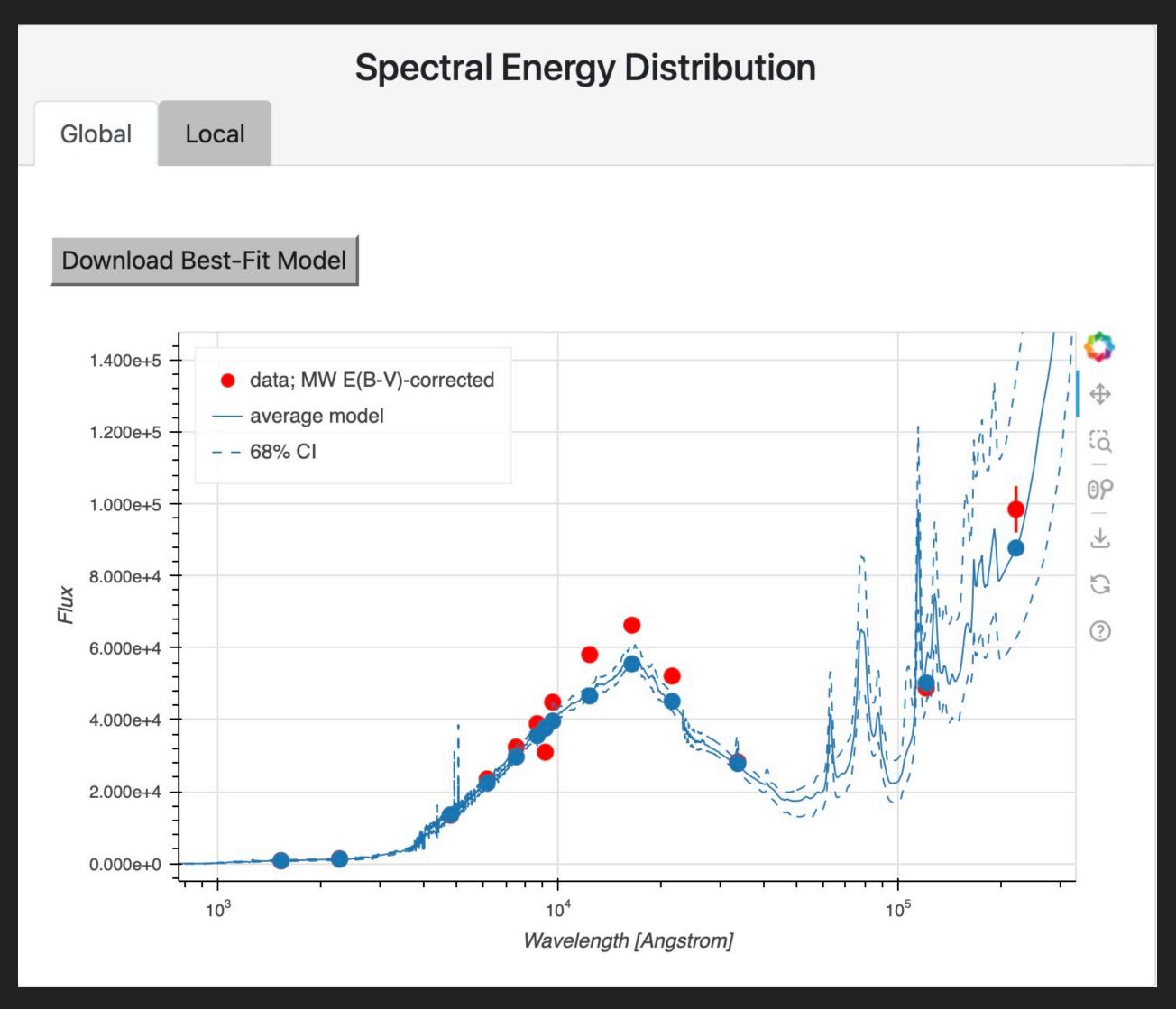
Local

Aperture details

RA, DEC (J2000) = 9h29m16.39s, -20d22m44.99s semi major axis, semi minor axis (arcsec) = 33.64, 32.86 angle (degrees, anti-clockwise from east) = -12.23

| Filter | Flux (μJy) | Flux error (μJy) | Mag | Mag error | Mag sys |
|-------------|-----------------|-----------------------|--------|--------------|------------|
| GALEX_FUV | 549.65 | 18.15 | 17.050 | 0.036 | AB |
| GALEX_NUV | 905.88 | 15.32 | 16.507 | 0.018 | AB |
| PanSTARRS_g | 11495.13 | 13.13 | 13.749 | 0.001 | AB |
| PanSTARRS_r | 21060.75 | 18.79 | 13.091 | 0.001 | AB |
| PanSTARRS_i | 29860.67 | 22.75 | 12.712 | 0.001 | AB |
| PanSTARRS_z | 36491.76 | 32.05 | 12.495 | 0.001 | AB |
| DES_z | 29218.61 | 80.14 | 12.736 | 0.003 | AB |
| PanSTARRS_y | 42539.06 | 65.52 | 12.328 | 0.002 | AB |
| 2MASS_J | 56137.9 | 855.6 | 11.132 | 0.017 | Vega |
| 2MASS_H | 64853.72 | 1061.07 | 10.503 | 0.018 | Vega |
| 2MASS_K | 51369.05 | 1732.01 | 10.285 | 0.037 | Vega |
| WISE_W1 | 28028.55 | 550.89 | 10.082 | 0.021 | Vega |
| WISE_W3 | 48681.76 | 1300.33 | 7.008 | 0.029 | Vega |
| WISE_W4 | 98412.99 | 6369.32 | 4.797 | 0.070 | Vega |

BLAST



Host SED inference Global Local Global parameter details **Documentation** Download Chains Download Percentiles Lower (16%) Median (50%) Upper (84%) **Parameter** $\log_{10}(M_*/M_{\odot})$ 10.497 10.594 10.701 $\log_{10}({ m SFR})$ -0.251 0.186 0.513 $\log_{10}({ m sSFR})$ -10.832 -10.384 -10.026 stellar age 6.962 7.993 11.079

Metallicity parameters AGN parameters

Lower (16%) Med

-1.437

-1.870

| Parameter | Lower (16%) Med | lian (50%) Upp | er (84%) |
|------------------------|-----------------|----------------|----------|
| $\log_{10}(f_{AGN})$ | -2.777 | -1.720 | -1.327 |
| $\log_{10}(au_{AGN})$ | 1.028 | 1.272 | 1.594 |

Dust parameters

Parameter

 $\log_{10}(Z_*/Z_{\odot})$

 $\log_{10}(Z_{gas}/Z_{\odot})$

| Parameter Lo | wer (16%) Med | ian (5 |
|-----------------------|---------------|--------|
| $	au_2$ | 0.479 | (|
| δ | -0.012 | (|
| $	au_1/	au_2$ | 0.441 | (|
| Q_{PAH} | 0.211 | (|
| U_{min} | 2.177 | 1 |
| $\log_{10}(\gamma_e)$ | -3.760 | -3 |

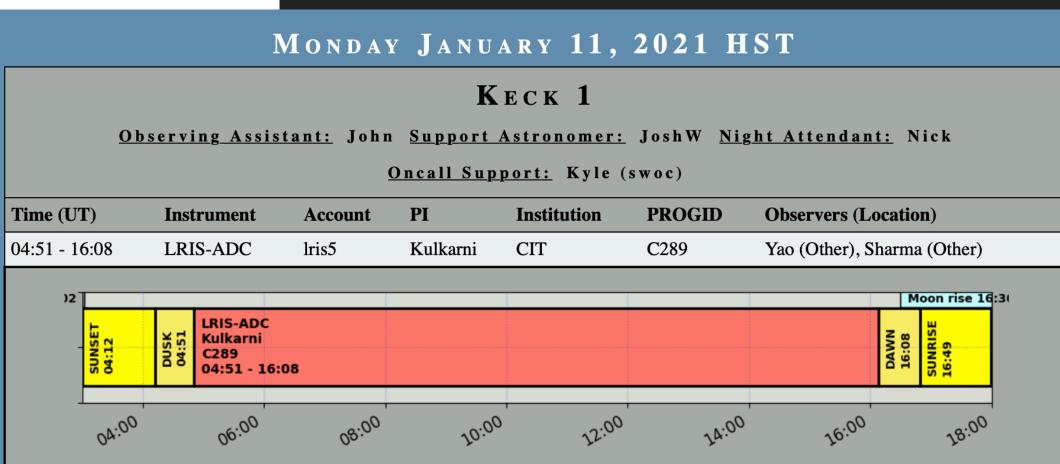
Binned star formation history

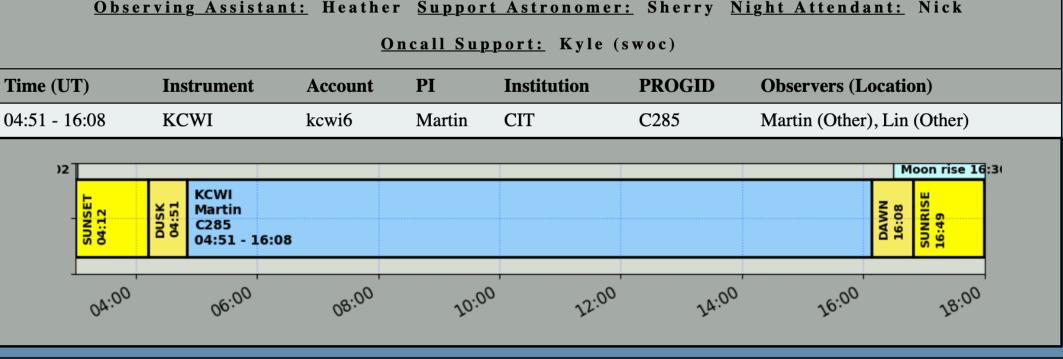
| Lookback time | Lower (16%) N | /ledian (50%) l | Jpper (84%) |
|-----------------|---------------|-----------------|-------------|
| 0.00-0.03 Gyr | -1.626 | -0.241 | 0.263 |
| 0.03-0.10 Gyr | -0.428 | 0.235 | 0.637 |
| 0.10-0.33 Gyr | -0.639 | 0.157 | 0.557 |
| 0.33-1.10 Gyr | -1.426 | -0.039 | 0.384 |
| 1.10-3.63 Gyr | -1.347 | 0.167 | 0.645 |
| 3.63-12.03 Gyr | 0.117 | 0.677 | 0.851 |
| 12.03-13.36 Gyr | 0.148 | 0.889 | 1.663 |

RESOURCE AVAILABILITY

- Example: you want to make an observation of a transient immediately. Which facilities can you use?
- What is my target's observability at each facility.
- This information has not been centralized it is distributed across hundreds of web sites with no common protocol.

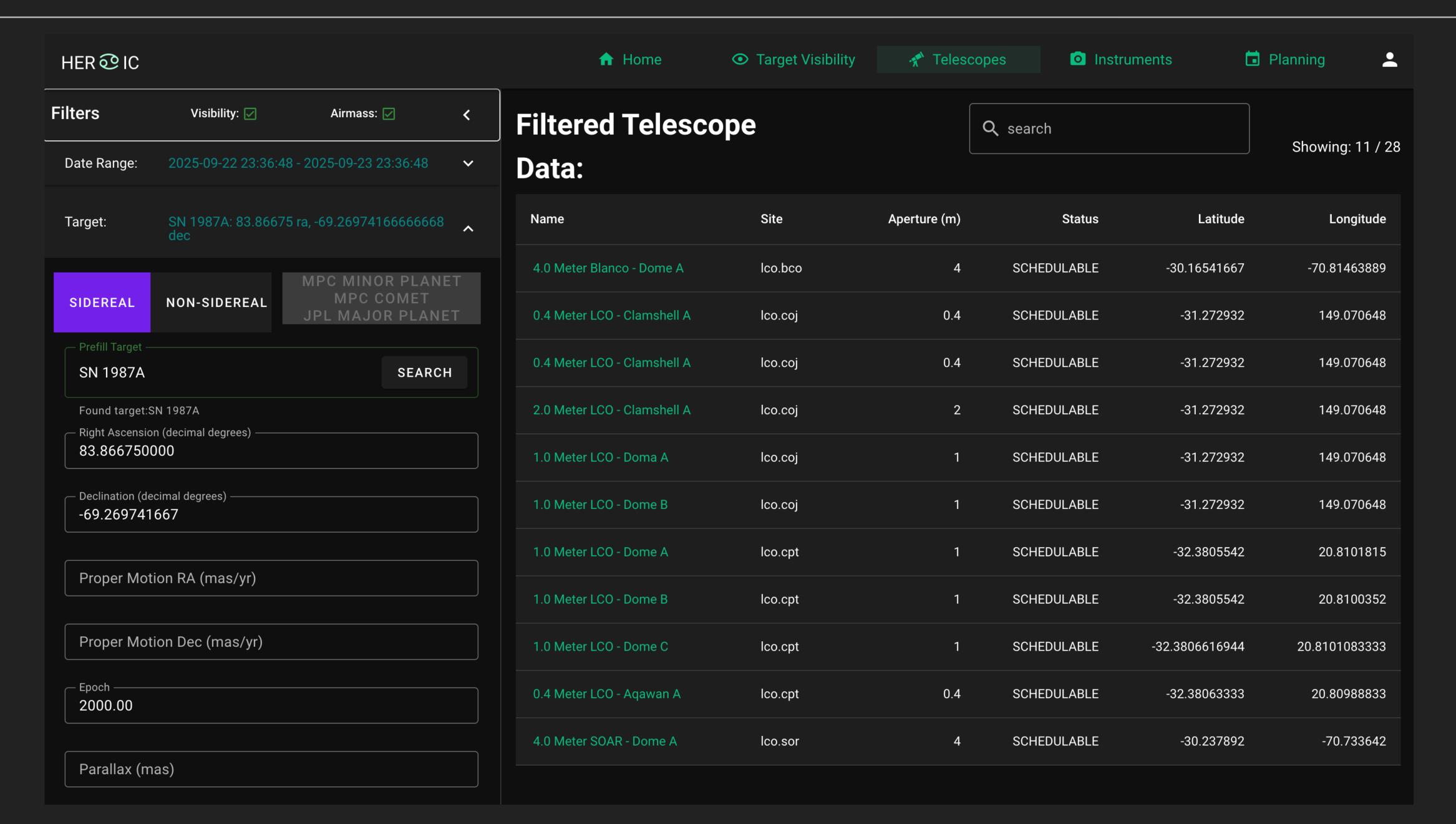
Mauna Kea at Mon Jan 11 16:01:06 2021 Gratings currently available in GMOS-N today... mirror B600+ G5307 R831+_G5302 R400+_G5305 Slits currently available in GMOS-N today... IFU-2 0.5arcsec 0.75arcsec 1.0arcsec 1.5arcsec 2.0arcsec 5.0arcsec NS0.75arcsec NS1.0arcsec 0.25arcsec



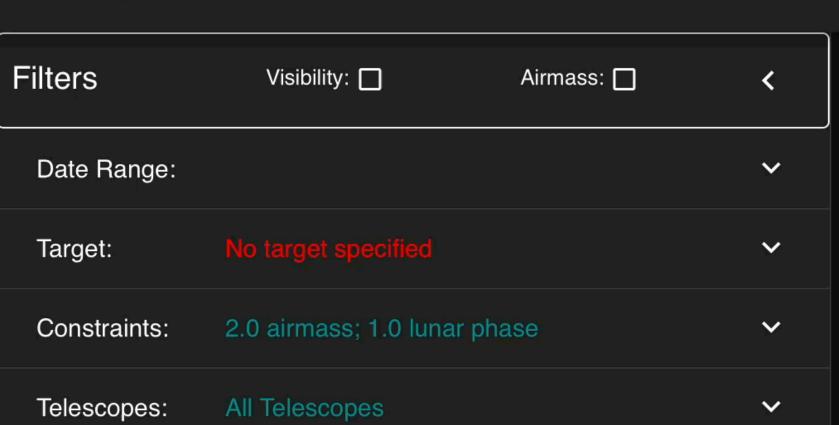


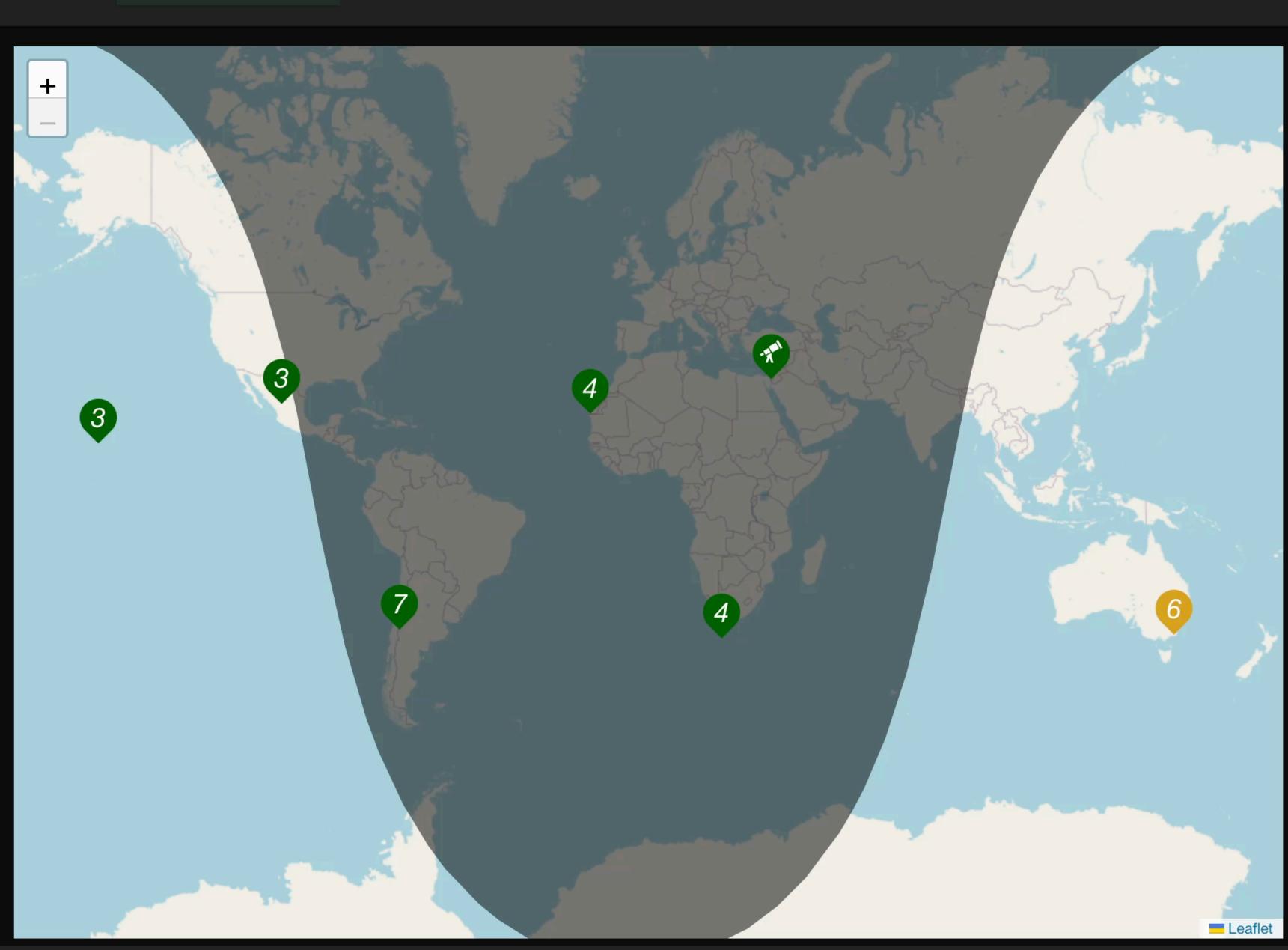
KECK 2

https://heroic.scimma.org



HER **②** IC





* Telescopes

★ Home

Target Visibility

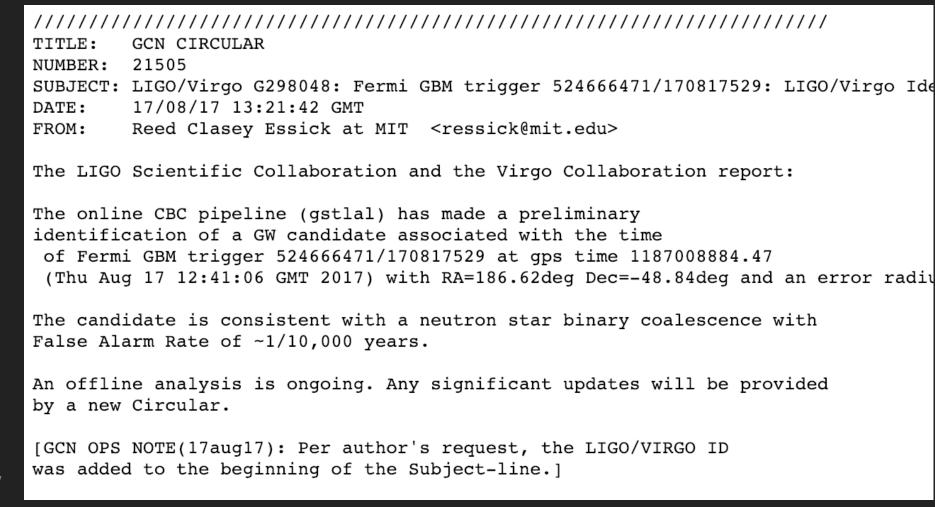
Instruments

☐ Planning

PREEXISTING MESSAGING SYSTEMS

GCN Circulars Used for GRBs & GW events. Free text, not machine readable.

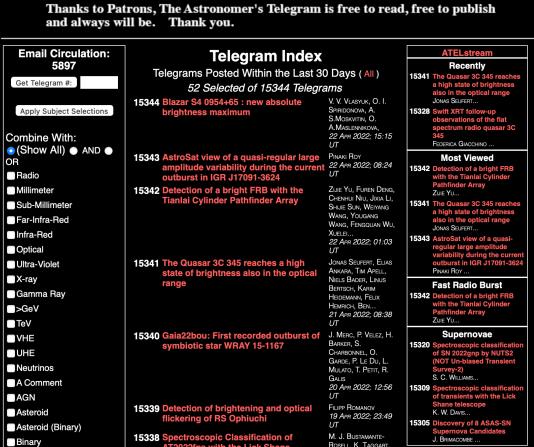
GCN Notices Machine readable, but only used by some NASA missions, no real free text.





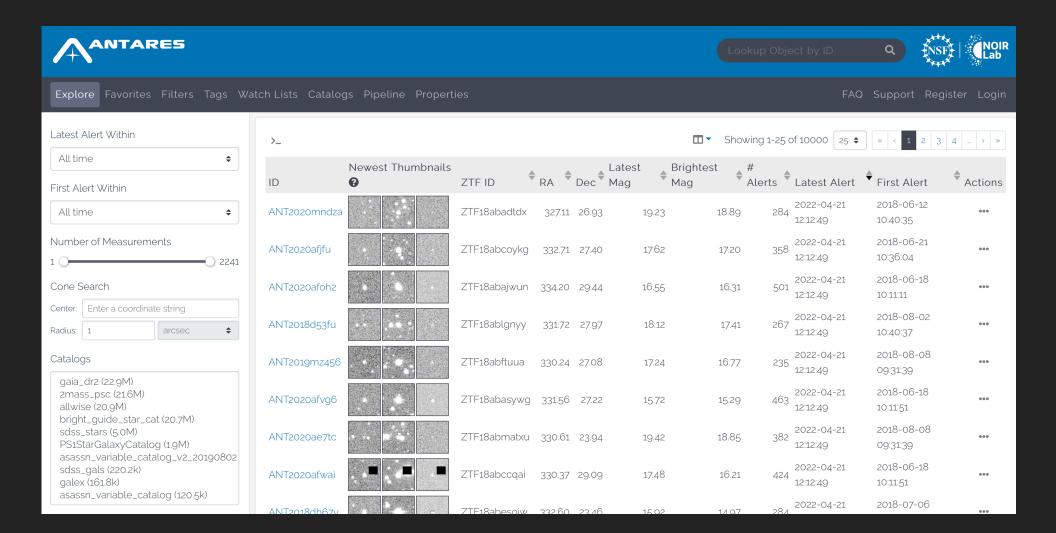


Thanks to Patrons, The Astronomer's Telegram is free to read, free to publish and always will be. Thank you.



Astronomer's **Telegrams (ATels)** Not machine readable. Largely abandoned by transient community.

information.



send classification spectra. Works well for specific, functions, but not used for sharing arbitrary

Brokers Designed for high alert volume, e.g. ZTF, LSST. Can use filters. No easy way as yet for users to report additional information.

HOPSKOTCH

- New messaging system being built by SCIMMA (Scalable Cyberinfrastructure for Multimessenger Astrophysics) funded by the NSF - see scimma.org
- Pub-sub model only subscribe to the information you want. This allows carrying much more information.
- Will carry existing existing astronomical messages, e.g. GCN Circulars and Notices, possibly Transient Name Server messages, Astronomer's Telegrams.
- Goal is to increase machine readable information.
- Can ultimately support other types of messages, e.g. sending images, spectra, data points, observation plans, instrument availability.



- Based on Kafka will scale to LSST era.
- Cloud-based. Hosted by Amazon Web Services
- Integrated with Identity and Access
 Management system (CILogon & KeyCloak).

DIFFERENCE BETWEEN HOPSKOTCH AND HERMES



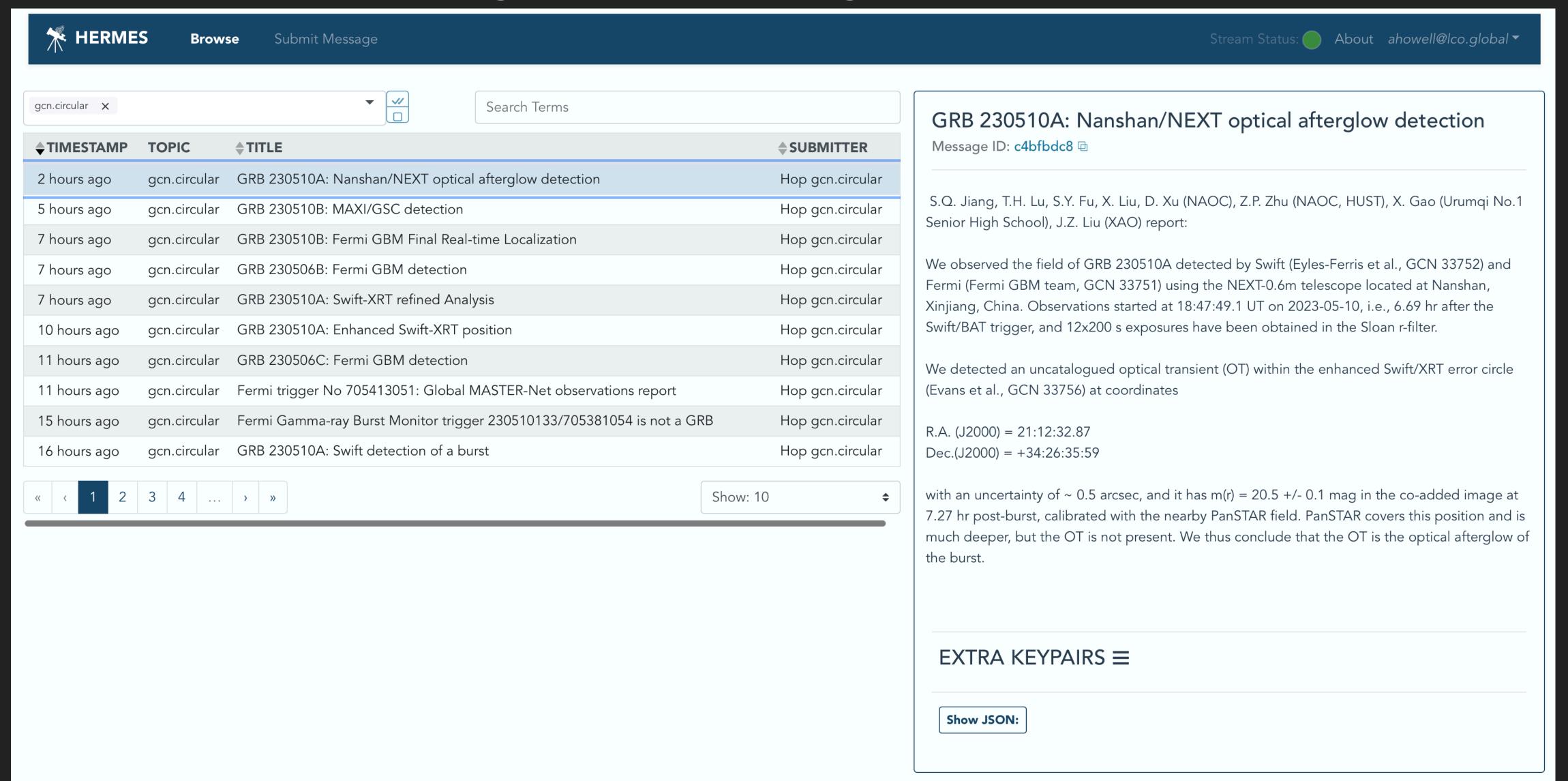
- Kafka + Libraries
- Identity and Access Management
- Topic creation, permissions
- Can carry any message type (e.g. GCNs, TNS, something new) – no schema
- Archive



- Built on HOPSKOTCH
- Reader and GUI
- API for submitting certain well-defined message types
- Machine-readable protocol

https://hermes.lco.global

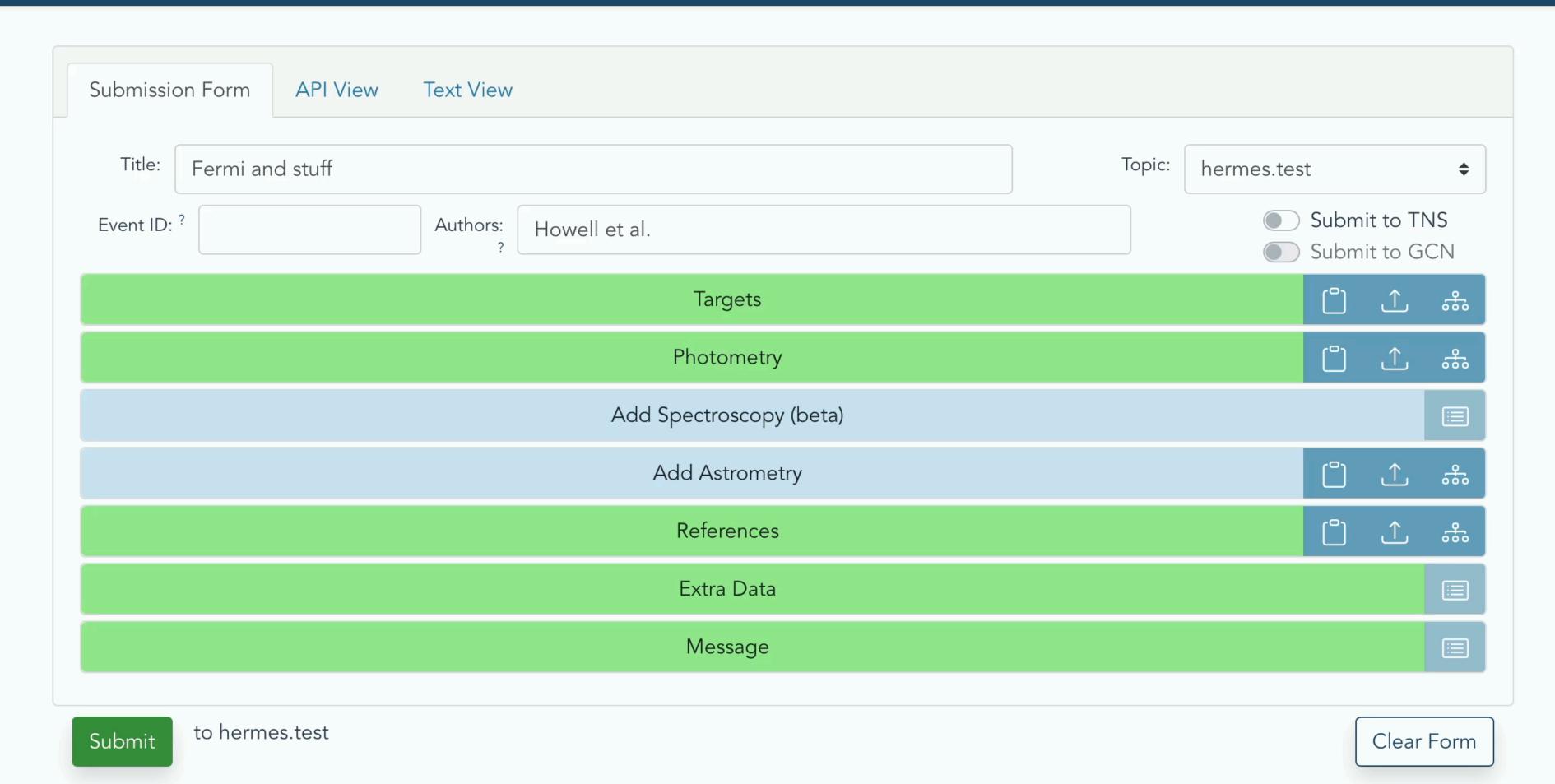
HERMES: Hopscotch Enabled Rapid Message Exchange Service Adds a message browser, GUI, API, message schema to HOPSKOTCH

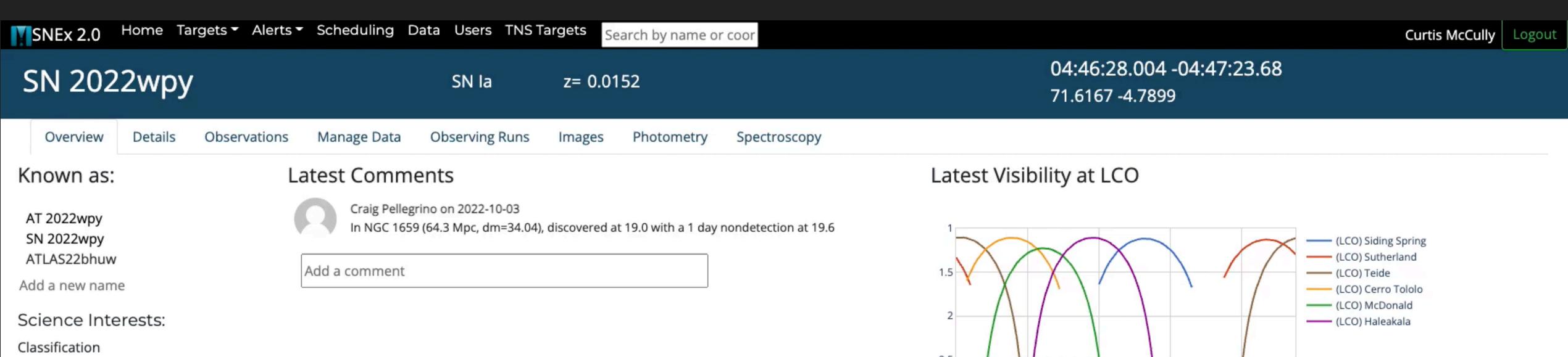


| Filter by Topic | | Search Terms | |
|-----------------|--------------|---|-----------------------|
| ♦TIMESTAMP | TOPIC | \$TITLE | ♦SUBMITTER |
| 11 minutes ago | igwn.gwalert | MS230511f - RETRACTION | cody.messick-30d512c8 |
| 16 minutes ago | igwn.gwalert | MS230511f - PRELIMINARY | cody.messick-30d512c8 |
| 21 minutes ago | igwn.gwalert | MS230511f - PRELIMINARY | cody.messick-30d512c8 |
| an hour ago | igwn.gwalert | MS230511e - RETRACTION | cody.messick-30d512c8 |
| an hour ago | igwn.gwalert | MS230511e - PRELIMINARY | cody.messick-30d512c8 |
| an hour ago | igwn.gwalert | MS230511e - PRELIMINARY | cody.messick-30d512c8 |
| 2 hours ago | igwn.gwalert | MS230511d - RETRACTION | cody.messick-30d512c8 |
| 2 hours ago | igwn.gwalert | MS230511d - PRELIMINARY | cody.messick-30d512c8 |
| 2 hours ago | igwn.gwalert | MS230511d - PRELIMINARY | cody.messick-30d512c8 |
| 3 hours ago | gcn.circular | GRB 230510A: Nanshan/NEXT optical afterglow detection | Hop gcn.circular |
| « | 3 4 > | » | Show: 10 |

HERMES is a Message Exchange Service for Multi-Messenger Astronomy applications that allow users to both send and review messages related to a variety of events and targets of interest.

Browse





2.5 Nearby SNe Young SNe Hours From Now Select Science Tags



J. Craig Wheeler Or Graur

I'm Interested

Add to Interesting Targets

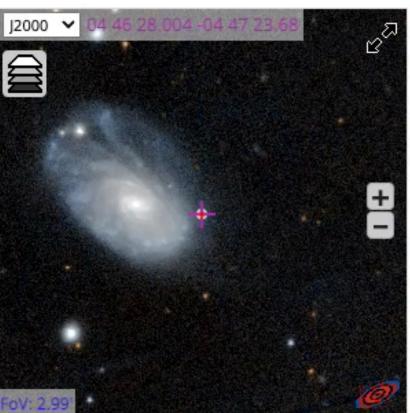
Data Used In:

First name of first author

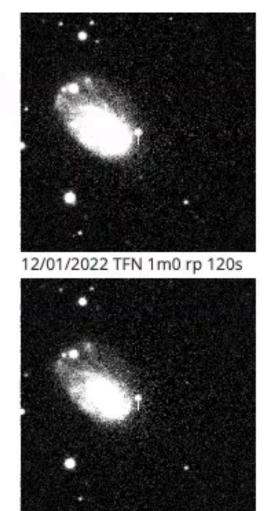
Last name of first author

Brief description of contents of this paper, i.e. "All photometry and

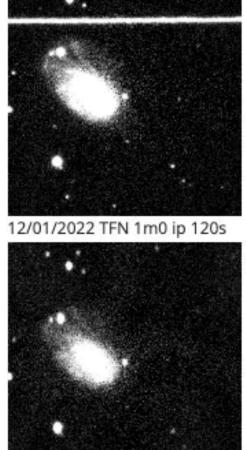
Aladin Viewer



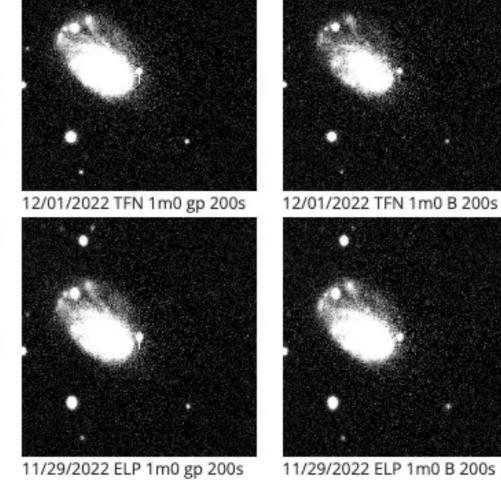
Recent LCO Images

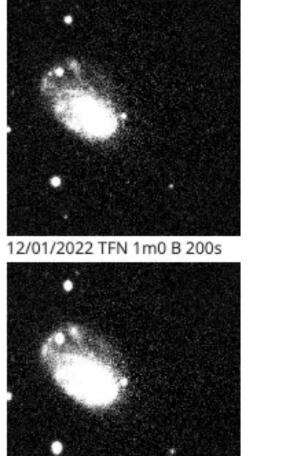


12/01/2022 TFN 1m0 V 120s



11/29/2022 ELP 1m0 ip 120s





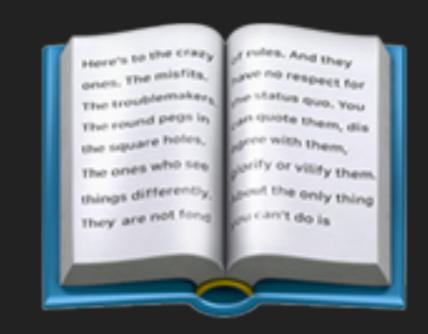
MODERN MESSAGING IS HERE



Low latency



Robust



Machine and human readable



Pub-sub: can distribute any message

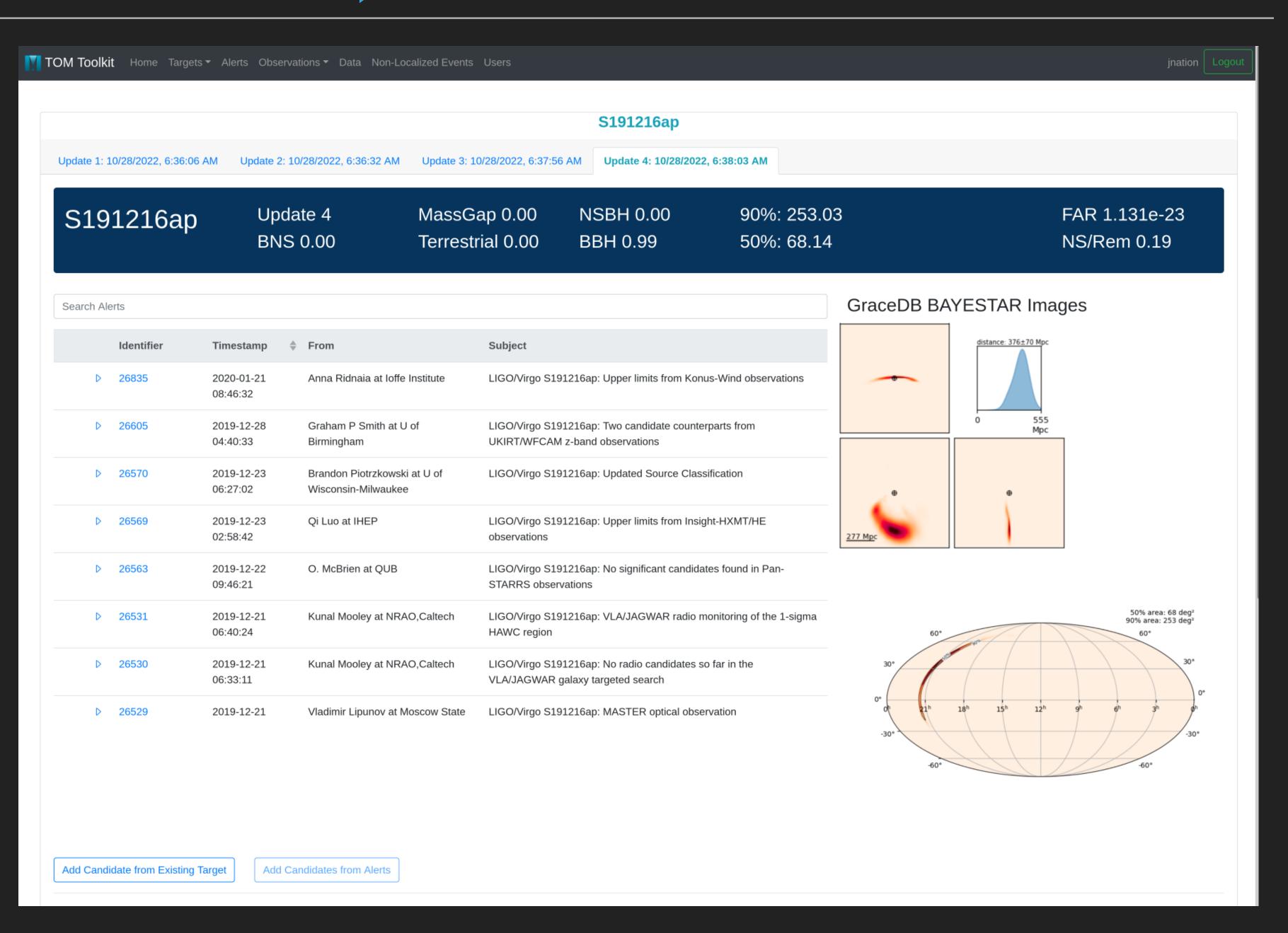
EXTRA SLIDES

| Filter by Topic | | Search Terms | |
|-------------------|--------------|---|-----------------------|
| ♦TIMESTAMP | TOPIC | \$TITLE | ♦SUBMITTER |
| 9 minutes ago | igwn.gwalert | MS230511f - RETRACTION | cody.messick-30d512c8 |
| 14 minutes ago | igwn.gwalert | MS230511f - PRELIMINARY | cody.messick-30d512c8 |
| 18 minutes ago | igwn.gwalert | MS230511f - PRELIMINARY | cody.messick-30d512c8 |
| an hour ago | igwn.gwalert | MS230511e - RETRACTION | cody.messick-30d512c8 |
| an hour ago | igwn.gwalert | MS230511e - PRELIMINARY | cody.messick-30d512c8 |
| an hour ago | igwn.gwalert | MS230511e - PRELIMINARY | cody.messick-30d512c8 |
| 2 hours ago | igwn.gwalert | MS230511d - RETRACTION | cody.messick-30d512c8 |
| 2 hours ago | igwn.gwalert | MS230511d - PRELIMINARY | cody.messick-30d512c8 |
| 2 hours ago | igwn.gwalert | MS230511d - PRELIMINARY | cody.messick-30d512c8 |
| 3 hours ago | gcn.circular | GRB 230510A: Nanshan/NEXT optical afterglow detection | Hop gcn.circular |
| « | 3 4 > | » | Show: 10 + |

HERMES is a Message Exchange Service for Multi-Messenger Astronomy applications that allow users to both send and review messages related to a variety of events and targets of interest.

NONLOCALIZED EVENT PAGE IN A TOM, CREATED FROM HERMES MESSAGE

- Nonlocalized event
 (e.g. coordinates
 not well known like
 in a GW error
 region) workflow
 added to TOM
 Toolkit
- Your TOM can monitor HERMES and create a page for a GW event automatically thanks to machine readability.
- Can keep track of viable and nonviable candidates





Search by name or coordinates

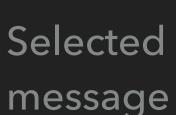
Authorship Information

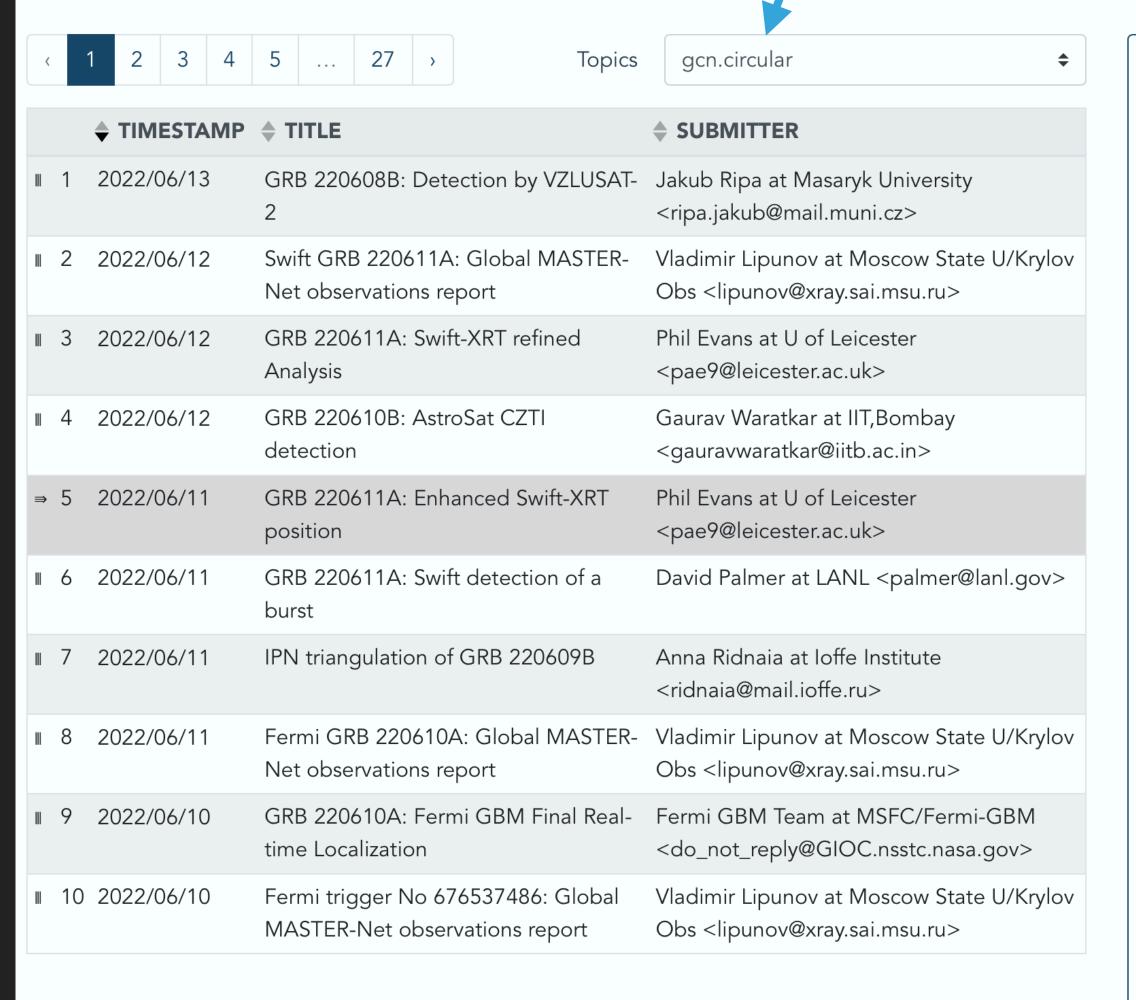
Latest Comments



Noshin Yesmin about at2023dtc on 2023-05-10 I ran SNID on both the first spectrum (2023-03-28) and on the latest (till today) (2023-04-26). For the first spectrum, top 10 type match was Type Ic and top 10 subtype match was Ic-norm. For the other spectrum, top 11 type

| Latest Targets | |
|----------------|------------|
| ID | Created |
| grb230506c | 2023-05-07 |
| grb230506b | 2023-05-07 |
| sn 2023bhz | 2023-05-05 |
| at 2023hnl | 2023-05-02 |
| sn 2023hlf | 2023-05-01 |
| at 2023gss | 2023-04-25 |
| at2023ghq | 2023-04-24 |
| at 2023gft | 2023-04-20 |
| ztf18aaxeoqk | 2023-04-20 |
| at2023gfo | 2023-04-20 |





HERMES Submit New Non-localized Event Submit Candidates

GRB 220611A: Enhanced Swift-XRT position

Phil Evans at U of Leicester <pae9@leicester.ac.uk>

ubmit Non-Detections Submit Search Pointings Report Photometry Report Spectroscopy Report Observatory Availability

M.R. Goad, J.P. Osborne, A.P. Beardmore and P.A. Evans (U. Leicester) report on behalf of the Swift-XRT team.

Using 496 s of XRT Photon Counting mode data and 1 UVOT images for GRB 220611A, we find an astrometrically corrected X-ray position (using the XRT-UVOT alignment and matching UVOT field sources to the USNO-B1 catalogue): RA, Dec = 66.51464, -37.26019 which is equivalent to:

RA (J2000): 04h 26m 3.51s Dec (J2000): -37d 15' 36.7"

with an uncertainty of 2.4 arcsec (radius, 90% confidence).

This position may be improved as more data are received. The latest position can be viewed at http://www.swift.ac.uk/xrt_positions. Position enhancement is described by Goad et al. (2007, A&A, 476, 1401) and Evans et al. (2009, MNRAS, 397, 1177).

This circular was automatically generated, and is an official product of the Swift-XRT team.

JSON / data

User▼

ADDITIONAL DATA TABLE ≡

Show JSON:

event.

Manually type or upload a data table.

There is built-in error checking.

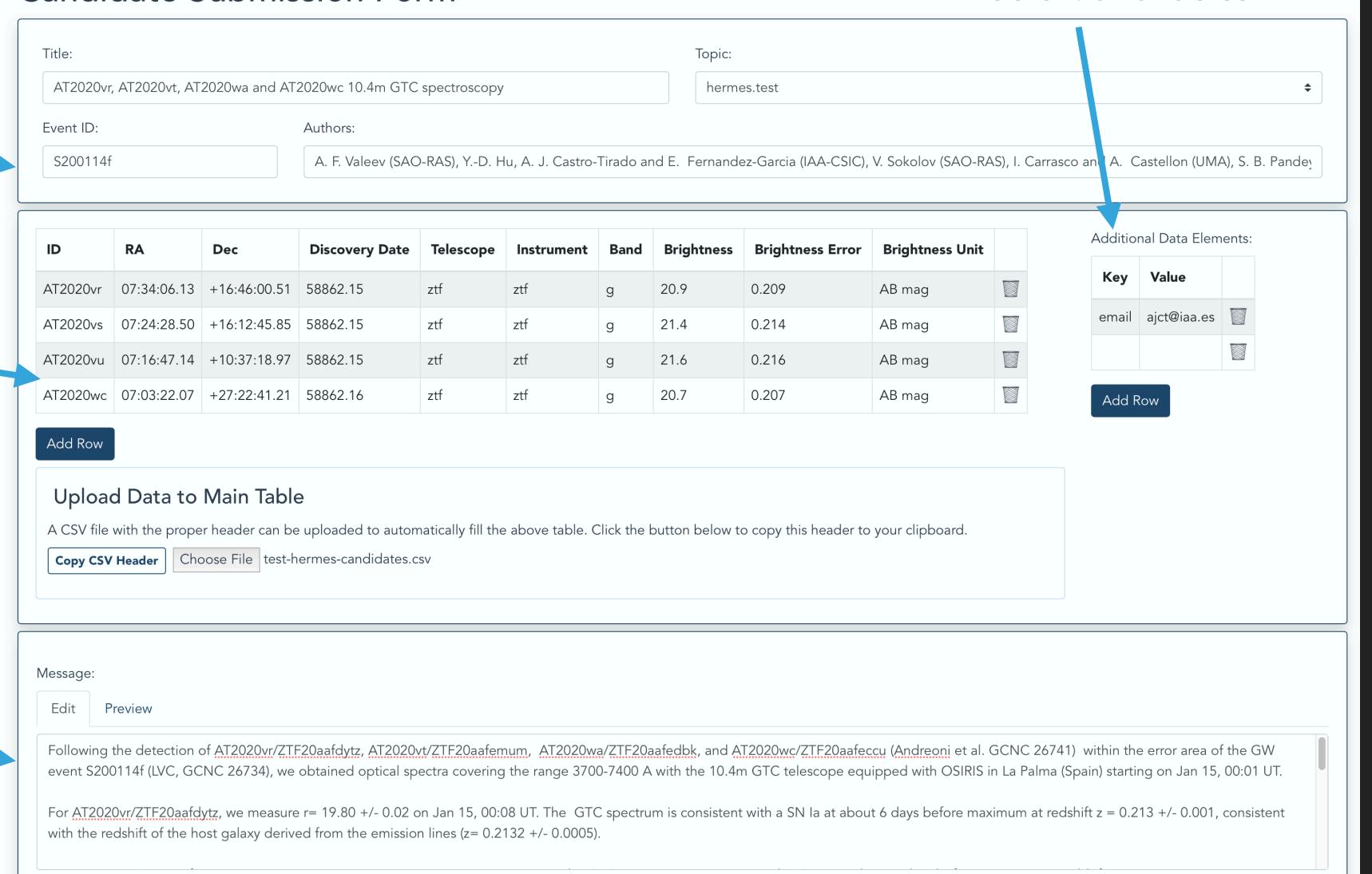
Compose a text message, and refer to items in the table.



HERMES Submit New Non-localized Event Submit Candidates Submit Non-Detections Submit Search Pointings Report Photometry Report Spectroscopy Report Observatory Availability

Candidate Submission Form

Add extra variables



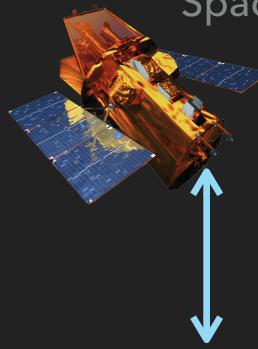
Submit

Clear Form









Multimessenger trigger

Broker













GUI + API + Protocol + Database

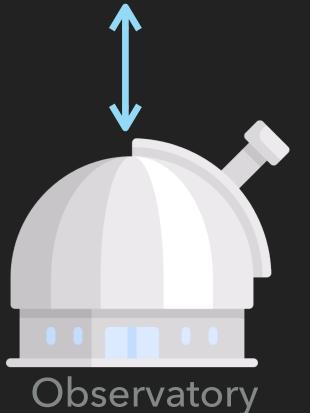
New Kafka tools



Coordinate, visualize follow-up



Target & Observation Manager



A NEW SOCIOLOGY?

- Right now we have either data that is fully public or fully private.
- This leads to duplication of data and inefficient use of resources because people don't know what others have found.
- We are disincentivized from sharing in real time because someone else can publish our data without us if we make it public.
- We need something in-between, a way for people to share preliminary data publicly, but still get credit.
- **Proposed solution:** a license when you share data that explains how it can be used, like software licenses. For example, "you can use this data for planning purposes and inferences, but if you want to publish it you have to ask me." This could be a flag sent with the data.

SUMMARY

- We need new communication infrastructure and software tools to fully realize the potential of LIGO-Virgo-Kagra, Vera Rubin Observatory, and other new facilities
- We are making improvements in nearly every step of this process, especially AEON for streamlining triggering, Treasure Map for finding targets, HEROIC for resource availability, the TOM Toolkit for data reduction, analysis, telescope triggering, and HOPSKOTCH & HERMES for messaging. Many of these were built to work together but can also work on their own.
- Some problems are still unsolved, especially data lakes and archive compatibility
- MMA is a useful driver because it has some of the most stringent requirements, but these tools should be useful for many other areas of transient astronomy.